

Dynamo action in a Radiatively Inefficient Accretion Flow

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Jets are observed in the low/hard state of a transient black hole binary (BHB). Jet production is thought to be connected to the presence of a radiatively inefficient accretion flow (RIAF) and a large-scale magnetic field close to the accreting black hole (BH).¹ However, the source of the large scale field is not quite obvious. It is possible to have an in-situ generation of the magnetic field in the accretion flow by a dynamo action. We investigate the magnetorotational instability (MRI)² driven dynamo in a geometrically thick ($H/R \sim 1$) RIAF by performing 3D global MHD simulations.³ We use the mean-field paradigm to describe the dynamo action in the RIAF.⁴ We observe an intermittent dynamo cycle (see Fig. 1) unlike the regular cycles seen in the simulations of a geometrically thin ($H/R \ll 1$) standard disc.

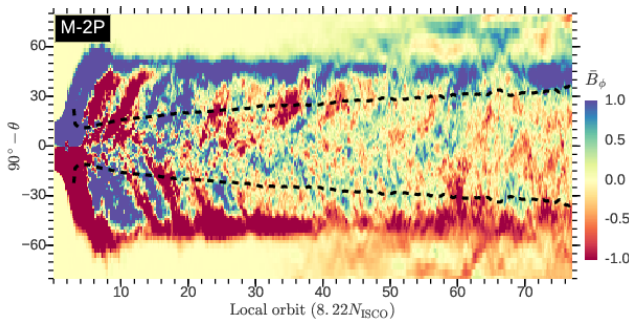


Fig. 1. Butterfly diagram: spatio (latitude)-temporal variation of mean toroidal magnetic field at a radius $r=20$. Time is expressed in units of local orbit at that radius.

Intermittency in the dynamo cycle is related to the slightly sub-Keplerian nature of the angular velocity.³

For the first time, we also obtain the spatial distribution of the turbulent dynamo coefficients in the poloidal plane (r, θ) for the MRI dynamo in the accretion flow.⁴ We find the evidence of a weak α -effect that is inefficient to produce ordered poloidal magnetic fields. Additionally, we observe the presence of a strong turbulent pumping transporting mean magnetic fields radially outward as shown in Fig. 2. We conclude that dynamo action is inefficient in generating ordered large-scale magnetic field in the RIAF. Assuming a truncated disc to exist in the low/hard state of BHBs, we propose a plausible scenario where the ordered large-scale

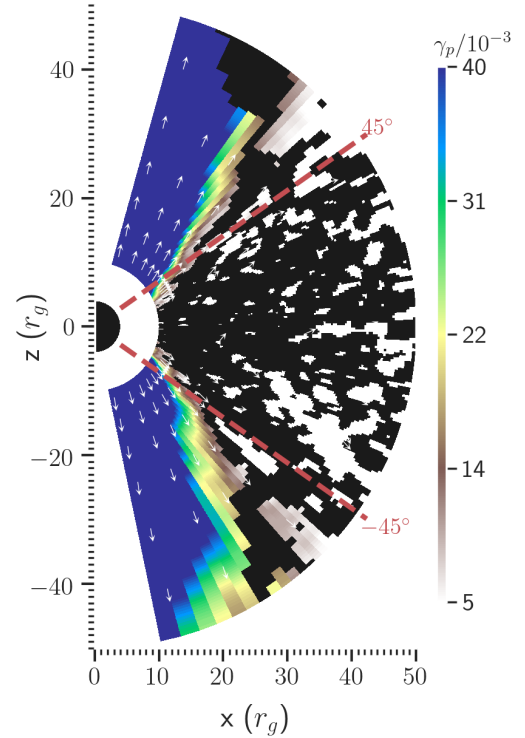


Fig. 2. Poloidal distribution of turbulent pumping term (combination of radial and meridional components). Colour shows the magnitude and arrows the direction. A black mask is applied in the regions where the ratio of the standard deviation to respective quantity's magnitude is greater than unity.

magnetic fields can be accumulated close the BH giving rise to a favourable condition for the production of the jets.

References

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